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Rapid Increase in Genetic Diversity in an Endemic Patagonian Tuco-Tuco Following a Recent Volcanic Eruption

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Rapid Increase in Genetic Diversity in an Endemic Patagonian Tuco-Tuco Following a Recent Volcanic Eruption

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1 **Rapid increase in genetic diversity in an endemic Patagonian tuco-tuco following a recent**
2 **volcanic eruption**

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20

21 Running header: Rapid genetic change in Patagonian rodents

22

23 Catastrophic natural events can have profound impacts on patterns of genetic diversity. Due to the
24 typically unpredictable nature of such phenomena, however, few studies have been able to
25 directly compare patterns of diversity before and after natural catastrophic events. Here, we
26 examine the impacts of a recent volcanic eruption in southern Chile on genetic variation in the
27 colonial tuco-tuco (*Ctenomys sociabilis*), a subterranean species of rodent endemic to the area
28 most affected by the June 2011 eruption of the Puyehue-Cordón Caulle volcanic complex. To
29 provide a comparative context for interpreting changes in genetic variation in this species, we
30 also analyze the effects of this eruption on genetic variation in the geographically proximate but
31 more widely distributed Patagonian tuco-tuco (*C. haigi*). Our analyses indicate that while both *C.*
32 *sociabilis* and *C. haigi* displayed significant post-eruption decreases in population density, the
33 apparent impacts of the eruption on genetic diversity differed between species. In particular,
34 genetic diversity at multiple microsatellite loci increased in *C. sociabilis* after the eruption while
35 no comparable post-eruption increase in *C. haigi* was observed at these loci. No changes in post-
36 eruption diversity at the mitochondrial cytochrome b locus were detected for either species. To
37 place these findings in a larger spatio-temporal context, we compared our results for *C. sociabilis*
38 to genetic data from additional modern and ancient populations of this species. These
39 comparisons, combined with Bayesian serial coalescent modeling, suggest that post-eruption
40 gene flow from nearby populations represents the most probable explanation for the apparent
41 increase in post-eruption microsatellite diversity in *C. sociabilis*. Thus, detailed comparisons of
42 pre- and post-eruption populations provide important insights into not only the genetic
43 consequences of a natural catastrophic event, but also the demographic processes by which these
44 changes in genetic diversity likely occurred.

45 Key words: *Ctenomys*, genetic diversity, subterranean rodents, volcanic eruption

46 Catastrophic natural events (e.g., hurricanes, floods, earthquakes) can severely impact
47 genetic diversity in natural populations, with such changes typically occurring within a short
48 period of time relative to the lifespan of the affected organisms (Pujolar et al. 2011). Such
49 catastrophic events often result in demographic bottlenecks that reduce genetic diversity and alter
50 the selective pressures acting on populations (Akey et al. 2004). However, determining the
51 precise demographic and genetic impacts of catastrophic events is challenging, in part because
52 the generally unpredictable nature of such events often precludes the collection of pre- and post-
53 catastrophic data, thereby necessitating inferences based solely on post-event sampling of
54 populations. Because few studies have directly compared pre- and post-event data, the specific
55 demographic processes underlying genetic responses to catastrophic change remain poorly
56 understood (Beheregaray et al. 2003; Pujolar et al. 2011; Wilmer et al. 2011).

57 The June 2011 eruption of the Puyehue-Cordón Caulle volcanic complex (40.5°S,
58 72.2°W, Fig. 1) in southeastern Chile provides a rare opportunity to explore the demographic and
59 associated genetic consequences of a naturally occurring catastrophic environmental event. The
60 eruption continued for more than 2 weeks (Collini et al. 2012), releasing more than 950 million
61 tons of tephra (ash and other organic matter) across northern Patagonia, from the Andean crest
62 eastward to the Atlantic Ocean (Gaitán et al. 2011). The impacts on livestock in the region were
63 substantial, with an estimated 60% decrease in populations of sheep and goats (Wilson et al.
64 2012). The eruption also impacted non-agricultural taxa; among exotic species, the eruption
65 caused fluoride intoxication and reduced longevity in European red deer (*Cervus elaphus*)
66 (Flueck and Smith-Flueck 2013; Flueck et al. 2014) as well as significant reductions in the
67 population sizes of introduced vespid wasps (Masciocchi et al. 2012). Native species were also
68 affected, as evidenced by reported reproductive abnormalities in liolaemid lizards (Boretto et al.

69 2014) and disruptions of both plant-pollinator relationships (Morales et al. 2014) and aquatic
70 invertebrate communities (Lallement et al. 2014). In sum, ash fall from the Puyehue-Cordón
71 Caulle eruption had widespread, diverse, and profound effects on the flora and fauna of northern
72 Patagonia.

73 Among the native taxa impacted by the ash fall were 2 species of tuco-tucos (Rodentia:
74 Ctenomyidae). Since 1992, the colonial tuco-tuco (*Ctenomys sociabilis*) and the Patagonian tuco-
75 tuco (*C. haigi*) have been the subjects of intensive field research aimed at characterizing the
76 behavior, ecology, and demography of each species (Lacey et al. 1997, 1998, 1999; Hadly et al.
77 2003; Chan et al. 2005; Lacey and Ebensperger 2007; Chan and Hadly 2011). As a result, data
78 from these herbivorous, burrow-dwelling taxa provide a rare opportunity to directly assess the
79 demographic and genetic consequences of the Puyehue-Cordón Caulle eruption. Both species
80 breed once per annum and have a generation time of 1 year (Chan et al. 2006). The colonial tuco-
81 tuco, listed as critically endangered by the IUCN, is endemic to a less than 1500 km² area in
82 southern Neuquen Province, Argentina, that includes the western side of the Limay River Valley
83 and adjacent hills (Fig. 1). In contrast, the parapatric *C. haigi* is widely distributed in the eastern
84 Limay Valley and adjacent portions of Río Negro Province, where it occurs in habitats similar to
85 those occupied by *C. sociabilis* (Lacey and Wieczorek 2004). Despite their close geographic
86 proximity and use of similar habitats, the 2 species differ markedly in their behavior: *C. haigi* is
87 solitary, with each adult occupying its own burrow system, while *C. sociabilis* is the only
88 ctenomyid species that has been demonstrated to live in groups (Lacey et al. 1997, 1998). The
89 long-term focal study populations for these species are located in the Limay Valley (Fig. 1),
90 approximately 100 km east of the Puyehue- Cordón Caulle volcanic complex. As a result, both
91 study sites experienced substantial (3-5 cm) ash fall following the 2011 eruption (Masciocchi et

92 al. 2012; Wilson et al. 2013) and both study populations experienced a significant post-eruption
93 decrease in population density (E. A. Lacey pers. obs.).

94 Here, we examine the impacts of the 2011 eruption on the genetic variation in the
95 populations of *C. sociabilis* that have been the focus of long-term demographic research.
96 Specifically, we compare levels of microsatellite and cytochrome b genetic variation in pre- and
97 post- eruption samples from these populations to assess the genetic consequences of this event.
98 To provide a comparative context for these analyses, we also analyze pre- and post-eruption
99 variation in the focal study population of *C. haigi*. To place our findings in a larger spatio-
100 temporal framework, we also consider pre-eruption data on genetic variation from several
101 additional populations of each species as well as available mitochondrial paleogenetic data from
102 these taxa, allowing us to contrast short-term changes in genetic diversity with more extended
103 responses to environmental change. Finally, we use Bayesian serial coalescent modeling
104 informed by demographic and genetic data from our study to identify population processes that
105 may have contributed to post-eruption changes in genetic diversity. Collectively, these analyses
106 offer important insights into the demographic processes underlying the immediate genetic
107 consequences of a natural catastrophic event.

108

109 **Materials and Methods**

110 *Sample collection and demographic analysis.*—Non-destructive tissue samples from *C. sociabilis*
111 and *C. haigi* were collected during the austral summers (November-January) of 1993-1998 and
112 2011-2013 (designated as pre- and post-eruption, respectively), following the trapping and
113 sampling procedures described in Lacey (2001). We analyzed samples from 85 individuals (54
114 pre-eruption, 31 post-eruption) from the focal population of *C. sociabilis* at Estancia Rincon

115 Grande and 62 individuals (45 pre-eruption, 17 post-eruption) from the focal population of *C.*
116 *haigi* at Estancia San Ramon (Table 1; Fig. 1). These localities have been the focus of an
117 intensive annual mark-recapture program since 1992 and thus genetic analyses of samples from
118 these populations could be linked to pre- and post-eruption demographic information from the
119 same animals (E. A. Lacey pers. obs.). To gain a more comprehensive picture of pre-eruption
120 genetic variation in these focal populations, we also included data from an additional 20
121 individuals per species from the pre-eruption time period; genotypes for these individuals were
122 generated as part of Lacey (2001). Given that sampling between the pre- and post-eruption
123 periods was separated by over a decade, none of the individuals included in the pre-eruption
124 sampling were present during or after the eruption. To assess the potential impacts of including
125 individuals of different ages in our analyses, we segregated the data set by age (adult versus
126 juvenile), ran separate analyses for each age class, and then compared the results from these
127 partitioned data sets to those from the full data set.

128 To place our comparisons of genetic variation in a broader spatial context, we included
129 samples from 4 additional populations (hereafter referred to as ancillary populations) of *C.*
130 *sociabilis* ($n = 42$ individuals) and 2 additional populations of *C. haigi* ($n = 25$ individuals)
131 collected between 1993 and 1998, during the pre-eruption time period (Fig. 1, Table 1). We
132 defined a population as a contiguous set of burrows that was spatially distinct from other
133 collections of burrows. These populations were clearly defined ecologically, given the distance
134 (~ 1 km) between clusters. The ancillary populations were not sampled after the eruption due to
135 difficulties accessing these sites under post-eruption conditions. Thus, direct comparisons of pre-
136 and post-eruption genetic variation were not possible for these localities. We have included data
137 from these ancillary localities, however, because they provide a broader spatial perspective on

138 pre-eruption genetic variation. The specific subsets of samples used in each analysis are indicated
139 in Table 1. Similarly, to provide a broader temporal context for our data, we included previously
140 published data on mitochondrial cytochrome b variation among 34 *C. sociabilis* and 31 *C. haigi*
141 (Chan et al. 2006) samples dating back to 12,000 years before present.

142 This study was carried out in compliance with all local, national, international, and
143 institutional regulations. Permits for fieldwork were granted by the Delegación Tecnica Regional
144 Patagonia de Parques Nacionales Argentinas and Provincia Rio Negro. All activities involving
145 live animals were approved by the Animal Care and Use Committee at the University of
146 California, Berkeley, and were consistent with the guidelines of the American Society of
147 Mammalogists for the use of wild mammals in research (Sikes et al. 2016).

148
149 *DNA extraction and microsatellite amplification.*— DNA was extracted from tissue samples
150 using a DNeasy Blood and Tissue extraction kit (Qiagen, Valencia, California). The success of
151 extractions was assessed via polymerase chain reaction, with negative controls included to test for
152 contamination. Following extraction, we amplified 5 *C. sociabilis*-specific microsatellite loci
153 using published primers from Lacey et al. (1999) and Lacey (2001). These loci were chosen
154 because 1) they represent the only species-specific loci known to be variable in *C. sociabilis*
155 (Lacey 2001), and 2) they were successfully genotyped in greater than 80% of our samples, thus
156 minimizing the potential for biases in the resulting data on genetic variation. Although other loci
157 have been genotyped for members of both species, previous analyses indicate that *C. sociabilis* is
158 monomorphic at these loci (Lacey et al. 1999; Lacey 2001), and thus analyses of these markers
159 were not expected to be informative regarding the effects of the 2011 eruption. The general

160 reaction protocol for PCR amplification and locus-specific annealing temperatures used were the
161 same as those specified in Lacey et al. (1999), Lacey (2001), and Chan et al. (2005).

162 *Microsatellite genotyping and scoring.*— Following PCR amplification of microsatellite loci, all
163 samples were sent to Molecular Cloning Laboratories (ABI 3730XL; South San Francisco,
164 California) or the Stanford Protein and Nucleic Acid facility (ABI 3130XL; Stanford, California)
165 for genotyping. The resulting fragment lengths were analyzed and individual genotypes were
166 assigned using Geneious Pro (Biomatters Limited). All genotypes were determined independently
167 by at least 3 different individuals; any discrepancies in allele assignments were resolved by re-
168 genotyping the locus for the sample in question. To further ensure accurate scoring of genotypes,
169 approximately 20% of DNA extracts from our post-eruption samples from *C. sociabilis* were re-
170 amplified or re-genotyped. In addition, multiple PCR products were sequenced by ElimBio
171 Pharmaceuticals (Hayward, California) to confirm the presence of a microsatellite repeat;
172 comparing these data against reference sequences in GenBank confirmed that flanking sequences
173 matched those from *C. sociabilis* and *C. haigi*. Finally, all samples were independently amplified,
174 genotyped, and scored by the Stanford Protein and Nucleic Acid facility to provide additional
175 verification; as a result, all samples in our data set were genotyped and scored in duplicate. The
176 resulting data were examined for evidence of null alleles and allelic dropout using MicroChecker
177 v2.2.3 (van Oosterhout et al. 2004). Based on these analyses (van Oosterhout et al. 2004), all loci
178 were retained for further analyses. We then pooled our data with genotypes for 20 pre-eruption
179 samples from each focal population obtained from Lacey (2001) to generate a more robust data
180 set for analyses of genetic variation. To ensure consistency across studies, a randomly chosen
181 subset of these samples analyzed by Lacey (2001) ($n = 6$) was re-genotyped to confirm allelic
182 assignments.

183 *Power analyses of microsatellite data.*— Given the relatively limited number of microsatellite
184 loci genotyped, we used POWSIM (Ryman and Palm 2006) to infer the statistical power of our
185 analyses. POWSIM employs a series of empirical parameters (number of loci, number of alleles
186 per locus, allele frequencies, and sample size), a specified number of populations, and a pre-
187 determined level of genetic diversity to estimate the power of a data set to detect genetic
188 differentiation. POWSIM analyses were run using a range of predefined F_{ST} values (0.01, 0.025,
189 and 0.05) to determine the power of our microsatellite data to detect potential differences in pre-
190 and post-eruption variation in each focal population as well as to detect differences between the
191 focal and ancillary populations of each study species.

192 In addition, to ensure that differences in sample sizes did not bias our comparisons of pre-
193 and post-eruption data, we conducted a bootstrapping analysis of microsatellite genotypes using a
194 custom R script (available upon request). For the focal population of each study species, the
195 number of pre-eruption samples analyzed (*C. sociabilis*, $n = 74$; *C. haigi*, $n = 65$) was greater than
196 the number of post-eruption samples (*C. sociabilis*, $n = 31$; *C. haigi*, $n = 17$). Accordingly, we
197 randomly sub-sampled 31 and 17 pre-eruption samples for *C. sociabilis* and *C. haigi*,
198 respectively. This sub-sampling was repeated 1,000 times, after which bootstrapped values were
199 compared to the observed values calculated for the entire data set to determine if estimates of
200 microsatellite variation were impacted by sample size.

201 *Analyses of microsatellite variation.*— To characterize microsatellite variation in the study
202 species, we tested for departures from Hardy-Weinberg equilibrium and calculated values for
203 standard population genetic parameters (observed and expected heterozygosity, F_{IS} , F_{ST} , allelic
204 richness and variance, genotypic differentiation, gene diversity) using GenePop (Raymond and
205 Rousset 1995), FSTAT (Goudet 1995), and Arlequin (Excoffier and Laval 2005). Allelic richness

206 was calculated per locus and population with a minimum sample size of 16 individuals. All
207 statistically significant outcomes for these and other analyses are explicitly indicated in the text of
208 the results or associated figures.

209 To evaluate spatial structuring of pre-eruption genetic diversity, we examined all
210 microsatellite data (pre- and post-eruption) from each focal population as well as data from the
211 associated ancillary populations for each species. Samples from focal and ancillary populations
212 had been collected during the same range of pre-eruption years (1993-1998). To explore potential
213 temporal genetic differentiation within the focal populations, we also included data from these
214 populations collected after the eruption (2011-2013). Assessments of potential spatial and
215 temporal genetic differentiation among conspecifics were conducted using the analysis of
216 molecular variance (AMOVA) test, as implemented in Arlequin. To determine if patterns of post-
217 eruption genetic variation in *C. sociabilis* and *C. haigi* revealed evidence of a recent decrease in
218 population size, we examined the microsatellite data for signatures of population bottlenecks
219 following the methods of Cornuet and Luikart (1996) and Garza and Williamson (2001).
220 Specifically, we tested for both heterozygosity excess and changes in the *M*-ratio, which
221 represents the ratio of the number of alleles to the range of allele sizes (Garza and Williamson
222 2001). Reductions in population size are expected to reduce the number of alleles more rapidly
223 than the range of allele sizes; thus, this ratio can reveal evidence of past bottlenecks (Garza and
224 Williamson 2001; Peery et al. 2012).

225 The use of 2 analytical methods that rely on different signals of change in a putatively
226 bottlenecked population provides additional power for detecting the effects of such historical
227 events. We used BOTTLENECK v1.2.02 (Piry 1999) to test for heterozygosity excess in each
228 post-eruption focal study population. We employed both a stepwise mutation model (S.M.M.)

229 and a 2-phase mutation model (T.P.M.) with the proportion of single mutations set at multiple
230 values (80%, 85%, 90%, 95%) considered appropriate for microsatellite loci (Funk et al. 2010).
231 We set the variance in mutation lengths at 0.36 and ran each test for 1,000 iterations. The
232 significance of any apparent excess in heterozygosity was assessed using a Wilcoxon signed rank
233 test. *M*-ratios were calculated with Arlequin, with critical values for this parameter determined
234 using *M*-crit (Garza and Williamson 2001). We set the average size of multi-step mutations to be
235 3.5 with 10% multi-step mutations in the model. Based on estimates provided by Chan et al.
236 (2006), we estimated the effective size for each focal population to be 300 individuals. We varied
237 the mutation rate from 10^{-3} to 10^{-4} , which corresponds to the range of mutation rates typically
238 found at microsatellite loci (Mapelli et al. 2012). Based on the equation $\theta = 4N_eM$, this thus
239 produced values of θ ranging from 1.2 to 12. Critical *M*-ratio values based on these values of
240 theta were then compared to the observed *M*-ratio values to determine the significance of
241 apparent signatures of bottlenecks.

242 *Mitochondrial DNA sequencing and analyses.*— To provide a potentially different perspective on
243 post-eruption changes in genetic variation and to enable comparisons of our pre- and post-
244 eruption samples with paleogenetic information from the same locus in the study species (Hadly
245 et al. 2003; Chan et al. 2005), we amplified portions of the mitochondrial cytochrome b locus (*C.*
246 *sociabilis*, $n = 900$ base pairs [bp]; *C. haigi*, $n = 600$ bp) across a subset of samples from both
247 species (*C. sociabilis*, $n = 45$; *C. haigi*, $n = 30$; Table 1). Amplification was performed using
248 primers MVZ05 (Smith 1998) and MVZ108 (Chan et al. 2005), which were designed to
249 encompass the regions of this locus that had been sequenced previously from paleogenetic
250 samples (Chan et al. 2006). PCR conditions followed those of Chan et al. (2005). Sequencing of
251 these products was completed by ElimBio Pharmaceuticals, after which the resulting sequences

252 were cleaned, assembled, and aligned using SeqMan (Lasergene Suite from DNAStar; Madison,
253 Wisconsin). Putative haplotypic variants were verified via visual inspection of the associated
254 electropherograms. All haplotypes generated were compared to cytochrome b sequences for *C.*
255 *sociabilis* and *C. haigi* available in GenBank. All novel sequences generated were accessioned to
256 GenBank (accession IDs KY013598-KY013609).

257 To characterize mitochondrial sequence diversity and examine potential signals of
258 demographic change in each species, we used Arlequin to calculate Tajima's D , Fu's F , F_{ST} and
259 F_{IS} values, θ_S (as defined by Watterson 1975), and θ_π (as defined by Tajima 1983) from our
260 cytochrome b sequences. Arlequin was also used to conduct an AMOVA for each species to
261 assess relative variation within versus among populations. To examine potential differences in
262 haplotype distributions over time, we generated haplotype network maps using TempNet (Prost
263 and Anderson 2011), which allows the temporal partitioning of heterochronous sequence data.
264 We generated a temporal haplotype network with the pre- and post-eruption samples from the
265 focal populations as distinct temporal layers. To provide a deeper temporal perspective on
266 haplotypic variation, we pooled cytochrome b sequence data generated by this study with
267 paleogenetic data from Hadly et al. (2003) and Chan et al. (2006) to examine potential changes in
268 haplotypic variation over the last 12,000 years. These analyses were completed using only those
269 portions of the modern cytochrome b sequence data (*C. sociabilis*, 398 bp; *C. haigi*, 288 bp,) that
270 corresponded to the paleogenetic sequences for each study species. Comparisons of modern and
271 paleogenetic data were conducted using TempNet.

272 *Bayesian serial coalescent modeling.*— To explore how potential demographic processes such as
273 changes in population size, mutation rate, and migration may have contributed to changes in
274 genetic diversity over time, we used BayeSSC (Anderson et al. 2005) with an approximate

275 Bayesian computation (ABC) framework to simulate the impacts of a bottleneck on the focal
276 study population of *C. sociabilis*. We used demographic estimates from trapping records (E. A.
277 Lacey pers. obs.) and from Chan et al. (2006) to simulate a 52.9% decrease in population size 2
278 generations prior to the collection of our data set; this timeline was chosen to encompass all
279 possible generations present at the time of our post-eruption sample. N_e prior to the bottleneck
280 was set at 50 individuals, a conservative estimate based on our demographic data. We simulated
281 the effects of the eruption for the microsatellite data, using a mutation rate of 0.001
282 mutations/generation. This figure represents the upper end of the range of mutation rates
283 considered biologically feasible for microsatellites (Mapelli et al. 2012). We ran simulations for
284 1,000,000 samples and calculated the percentage of runs that displayed a change in
285 heterozygosity matching the change observed post-eruption in our empirical data set.

286 We then explored the effects of changes in population size, mutation rate, and migration
287 rate on post-eruption genetic diversity. For the microsatellite data set, we varied the prior
288 distributions for these parameters in our simulations, which were again run for 1,000,000
289 samples. We applied an ABC framework with a 5% rejection threshold for the posterior
290 distributions to determine the most likely estimates for the priors. Specifically, we compared 4
291 summary statistics (pre- and post-eruption observed heterozygosities and allelic variances) from
292 the empirical data set to our posterior distributions to determine the most likely estimate (MLE)
293 for each prior. We ran analyses using the average observed heterozygosity and allelic variance
294 calculated across all microsatellite loci as well as the observed heterozygosity and allelic variance
295 from the 2 loci (Sociabilis 4 and Sociabilis 7) that had the largest and smallest post-eruption
296 changes in heterozygosity, respectively (see results). Use of these extremes provided an upper
297 and lower bound for our MLEs of priors. We varied the uniform distributions used for priors to

298 assess the consistency of the resulting posterior distributions and MLE values. In the event of
299 large variations in MLE values, we assessed the fit of standard statistical distributions (uniform,
300 exponential, normal, and gamma distributions) to the posterior using negative log likelihood and
301 Akaike Information Criteria (AIC) values and, if necessary, modified the associated prior.

302 The first parameter examined was population size, which was set as a uniform prior
303 ranging from 0 to 5,000 individuals, a conservative range chosen to encompass any biologically
304 feasible population size for this species. Next, to explore the impact of mutation rate, we ran
305 simulations with the mutation rate set as a prior and with a constant population size (pre-eruption
306 estimate of 50 individuals); for these runs, we used a uniform distribution of mutation rates
307 ranging from 0 to 0.3, which encompasses the range of substitution rates reported for mammalian
308 microsatellites (Yue et al. 2002; Mapelli et al. 2012). Finally, to examine the impacts of
309 migration, we repeated these simulations but allowed for the presence of a second population,
310 representing the pooled ancillary populations for the species. For these analyses, we set the
311 forward-in-time migration rate from the neighboring localities into the focal population of *C.*
312 *sociabilis* as a prior with a uniform distribution. We varied 1) N_e for the neighboring localities
313 (range = 50 to 10,000 individuals), and 2) forward-in-time migration rate from the focal
314 population to neighboring localities (migration rate ranged from 0 to 0.1). All simulations were
315 run for 100,000 samples. While simulations were run backward in time, all migration rates here
316 are presented as forward-in-time for clarity.

317

318 **Results**

319 Microsatellite genotypes were largely consistent across multiple independent rounds of
320 amplification, genotyping, and scoring, providing verification of our results (individual

321 microsatellite genotypes are available by request). Within species, we found no differences in
322 microsatellite variation between data sets containing adults only, juveniles only, or adults and
323 juveniles together and thus all subsequent analyses were conducted using all samples for each
324 species. Our bootstrapping analyses provided evidence that sample size did not affect estimates
325 of variation: with the exception of 1 locus in *C. haigi* (Sociabilis 1), all measures of
326 heterozygosity in our pre-eruption samples fell within 95% confidence intervals for estimates of
327 heterozygosity obtained via bootstrapping, suggesting that differences in sample sizes for pre-
328 and post-eruption data sets did not impact estimates of genetic diversity. The 1 locus that differed
329 in *C. haigi* (Sociabilis 1) had a mean bootstrapped heterozygosity value that was less than 0.01
330 different from the empirical value of heterozygosity across all samples, and thus was still
331 included in analyses to obtain more robust results. Accordingly, we included all available pre-
332 and post-eruption genotypes in our analyses of microsatellite variation.

333 Our POWSIM results indicated that our microsatellite data set had sufficient power to
334 detect genetic differentiation. A power of 0.80 or greater has been suggested as the threshold for
335 providing adequate statistical power (Cohen 1988); we found that the power of our comparisons
336 between pre- and post-eruption variation in our focal populations as well as between the focal and
337 ancillary populations in each species were largely greater than this threshold. Specifically, given a
338 predefined F_{ST} value of 0.025 or greater, the POWSIM analyses indicated that the data set for *C.*
339 *sociabilis* had a power of 0.75 to detect genetic differentiation (all values provided from chi-
340 squared test). Statistical power rose to 0.91 if these analyses were run with $F_{ST} > 0.05$. Estimated
341 power of the data set from the focal population of *C. haigi* was greater, with a power of 0.99 for
342 $F_{ST} > 0.025$ and a power of 1.00 for F_{ST} value > 0.05 . When the ancillary populations for each
343 species were included, power was 0.93 and 1.00 for *C. sociabilis* and *C. haigi*, respectively, for

344 $F_{ST} > 0.025$. Power rose to 0.99 and remained at 1.00 for *C. sociabilis* and *C. haigi*, respectively,
345 for $F_{ST} > 0.05$. Given that observed values of F_{ST} (see results section below) for both study
346 species were greater than those used in the POWSIM analyses, and given that the resulting
347 estimates of power were generally greater than the 0.80 threshold, our data set appeared to be
348 sufficient to detect genetic differentiation in both study species. Thus, we retained all
349 microsatellite loci in our analyses and used these loci to compare temporal and genetic
350 divergence among populations.

351 *Temporal changes in cytochrome b diversity.*— We analyzed cytochrome b sequences for 45
352 individuals from the focal population of *C. sociabilis* and 30 individuals from the focal
353 population of *C. haigi* (Table 1); this sample included data from 14 *C. sociabilis* and 13 *C. haigi*
354 sampled during the pre-eruption time period (1993-1998) that had been sequenced previously by
355 Lacey (2001) and Hadly et al. (2003). Only a single cytochrome b haplotype was detected in the
356 focal population of *C. sociabilis* both before (1993-1998, $n = 17$) and after (2011-2013, $n = 26$)
357 the 2011 eruption. In contrast, the focal population of *C. haigi* was characterized by multiple
358 haplotypes in both time periods (5 haplotypes pre-eruption; 10 haplotypes post-eruption; Fig. 2).
359 Expansion of our analyses to include cytochrome b genotypes from paleogenetic samples from
360 each species (Chan and Hadly 2011) revealed a general pattern of declining haplotype diversity in
361 *C. sociabilis* over the past 10,000 years (Supplementary Data S1A); in comparison, *C. haigi* has
362 maintained relatively consistent levels of haplotype diversity over this same time period
363 (Supplementary Data S1B).

364 Patterns of nucleotide diversity at the cytochrome b locus paralleled interspecific
365 differences in haplotypic diversity. With only 1 haplotype detected in both pre- and post-eruption
366 samples, there was no change in diversity in *C. sociabilis*. In contrast, analyses of haplotypes

367 from the focal population of *C. haigi* revealed considerably greater levels of nucleotide diversity
368 for all measures considered. Prior to the eruption, θ_S was 1.66 and θ_π was 1.24; post-eruption,
369 these values were 3.01 and 2.88, respectively. Values for Tajima's D (pre-eruption = -0.92, $P =$
370 0.22; post-eruption = -0.17, $P = 0.47$) were negative, as were values for Fu's F_s (pre-eruption = -
371 18.1; post-eruption = -18.55; P for both < 0.005). AMOVA analyses revealed that 97.8% of the
372 variation detected occurred within the study population, with the remainder occurring between
373 the pre- and post-eruption subsets of this population. The F_{ST} between pre- and post-eruption
374 samples was 0.0218 ($P = 0.25$). In sum, these analyses indicate that cytochrome b haplotypes
375 were characterized by considerably more nucleotide diversity in *C. haigi* than in *C. sociabilis*.
376 *Temporal changes in microsatellite diversity.*—In *C. sociabilis*, only 1 microsatellite locus
377 (Sociabilis 4) deviated significantly from Hardy-Weinberg expectations (Hardy-Weinberg
378 probability tests, 1000 iterations; $P < 0.005$; Table 2a) in the pre-eruption focal population,
379 although there was evidence of a significant departure from neutral expectations when data from
380 all loci were considered together (Fisher's exact test, $P < 0.05$). After the 2011 eruption, no loci
381 deviated from Hardy-Weinberg expectations in this population (Table 2a), and similarly there
382 was no deviation from Hardy-Weinberg expectations when data from all loci were considered
383 together (Fisher's exact test, $P = 0.18$). Prior to the eruption, 4 of 5 microsatellite loci displayed
384 positive values of F_{IS} ; in contrast, all post-eruption values of F_{IS} were negative (Table 2a; all $P >$
385 0.05).

386 The number of alleles detected in the focal population of *C. sociabilis* did not change
387 before versus after the eruption for 2 of 5 loci (Sociabilis 1 and 5) and decreased by only a single
388 allele after the eruption at each of the remaining loci (Sociabilis 4, 6, and 7). Similarly, pre- and
389 post-eruption allelic richness did not change appreciably for Sociabilis 1, 5, and 7 and decreased

390 only modestly for Sociabilis 4 and 6 (Table 2a). Comparisons of pre- and post-eruption genotypes
391 revealed significant temporal differentiation at only one locus (Sociabilis 1, exact G test, $P <$
392 0.05). In contrast, observed heterozygosity increased significantly after the eruption (Fisher's
393 combined probability test, $P < 0.05$); this pattern was evident for all loci genotyped and
394 significant for 2 loci (Fig. 3A). Values of gene diversity for all loci were greater after the
395 eruption, with mean diversity across all loci increasing from 0.23 to 0.32 (Table 2a). Similarly,
396 estimates of θ_H increased after the eruption for all loci. Although AMOVA analyses revealed that
397 only 4.53% of microsatellite variation was due to temporal differences among samples,
398 comparisons of pre- and post-eruption data sets indicated small but statistically significant
399 temporal population subdivision ($F_{ST} = 0.055$, $P < 0.05$). Thus, these analyses indicate that
400 microsatellite heterozygosity and most associated measures of genetic diversity in the focal
401 population of *C. sociabilis* were greater after the eruption.

402 Applications of the same analyses to microsatellite data from the focal population of *C.*
403 *haigi* revealed a substantially different pattern of pre- versus post-eruption genetic diversity. Prior
404 to the eruption, none of the 5 loci examined revealed significant departures from Hardy-Weinberg
405 expectations; post-eruption, only 1 locus (Sociabilis 6) deviated from Hardy-Weinberg
406 expectations (Table 2b). Across all loci, no evidence of significant departure from Hardy-
407 Weinberg expectations was found for either pre- or post-eruption data sets (Fisher's Exact Test,
408 pre-eruption $P = 0.63$; post-eruption $P = 0.48$). During both time periods, 2 of 5 loci were
409 characterized by negative F_{IS} values, although the identities of these loci differed for pre- versus
410 post-eruption data sets (Table 2b).

411 The number of alleles detected in the focal population of *C. haigi* decreased after the
412 eruption for 4 of the 5 loci examined; differences in allelic richness were generally small pre-

413 versus post-eruption, with the exceptions of loci Sociabilis 6 and 7 (Table 2b). There was no
414 consistent pattern of change in gene diversity across time periods, with the largest changes
415 occurring at 3 loci (Sociabilis 5, 6, and 7; Table 2b). Although observed heterozygosity did not
416 differ consistently or significantly before versus after the eruption (Fig. 3b), genotypic diversity at
417 each locus was reduced in the post-eruption samples (Fisher's Exact G test, all $P < 0.05$).

418 AMOVA analyses indicated that only 2.39% of microsatellite variation was due to temporal
419 differences among samples. Although temporal differences in allelic frequencies estimated by
420 F_{ST} were low, the difference in this statistic was significant ($F_{ST}=0.024$, $P < 0.05$). Thus, while
421 *C. sociabilis* was characterized by significant temporal partitioning of microsatellite variation and
422 generally greater microsatellite diversity after the eruption, *C. haigi* displayed no evidence of
423 strong temporal partitioning or greater microsatellite diversity after the eruption.

424 *Tests for population bottlenecks.*— A significant excess of heterozygosity was detected for the
425 post-eruption samples of *C. sociabilis* (Wilcoxon signed-rank test; 2-phase model with 80%
426 proportion of single mutations; $P < 0.05$). In contrast, no such excess was detected for the post-
427 eruption population of *C. haigi* (Wilcoxon signed-rank test; 2-phase model with 80% proportion
428 of single mutations; $P = 0.41$). These patterns were consistent when we varied the percentage of
429 single mutations allowed as well as when we repeated analyses using a single mutation model (*C.*
430 *sociabilis*, $P < 0.05$; *C. haigi*, $P = 0.92$). *M*-ratio tests (Cornuet and Luikart 1996) revealed that
431 for *C. sociabilis*, observed *M*-ratios for all loci were below critical *M* values for all estimates of θ
432 employed (Table 3). In contrast, for *C. haigi*, only 1 locus (Sociabilis 5) was characterized by an
433 *M*-ratio below the associated critical value. Collectively, these analyses indicate that the 2 focal
434 populations exhibited different genetic signals of recent demographic history, with only *C.*
435 *sociabilis* displaying patterns of microsatellite variation consistent with a recent population

436 bottleneck. *Modeling of demographic parameters.*— Serial Bayesian coalescent modeling based
437 on demographic data from *C. sociabilis* revealed that observed patterns of pre- and post-eruption
438 microsatellite variation differed between our empirical data and expected outcomes based on our
439 demographic models ($P < 0.05$). This finding suggests that the observed post-eruption increase in
440 microsatellite diversity in *C. sociabilis* was contrary to predictions based on the demography of
441 this species.

442 Application of an ABC modeling framework to our simulated microsatellite distributions
443 revealed that the most likely estimates (MLEs) of N_e for the focal population of *C. sociabilis*
444 ranged from 244 to 254 (Fig. 4A); these values are 4 to 5 times greater than empirical estimates
445 of N_e for this population (E. A. Lacey, pers. obs.), suggesting that the increase in N_e required to
446 generate the observed post-eruption increase in microsatellite diversity is biologically unlikely.
447 The MLEs for microsatellite mutation rate ranged from 0.0153 to 0.0165 (Fig. 4b); again, these
448 rates fall beyond the range of mutation rates considered biologically likely for microsatellites (10^{-3}
449 to 10^{-4} ; Mapelli et al. 2012), suggesting that mutation cannot account for the post-eruption
450 increase in diversity observed in *C. sociabilis*. When data from the pooled ancillary populations
451 were included in these simulations, posterior distributions revealed a better fit (sharp peaks in
452 posterior distributions, lower delta scores) when migration from the focal population to other
453 populations was allowed (for clarity, all migration rates are presented as going forward in time,
454 although simulations were run backward in time). The fit of posterior distributions increased as
455 the rate of migration from the focal population to other populations was increased or N_e for the
456 pooled additional populations was increased. Indeed, even a modest migration rate of 0.01 (1%
457 probability per generation for each individual migrating from the focal to the pooled ancillary
458 population), along with a small effective size ($N_e = 50$) for the pooled ancillary population

459 produced a peak indicating a MLE of 0.075 for migration into the focal population. Overall,
460 MLEs for migration rate into the focal population ranged from 0.0176 to 0.075 (Fig. 5), which
461 represents a maximum migration of 4.5 individuals per generation, a number that is biologically
462 plausible for this species (Lacey and Wiczorek 2004).

463

464 **Discussion**

465 Our analyses indicate that the 2011 eruption of the Puyehue-Cordón Caulle volcanic
466 complex in southern Chile impacted genetic variation in *C. sociabilis* and that the consequences
467 of this event differed between the group-living *C. sociabilis* and the parapatric, solitary *C. haigi*.
468 Long-term behavioral and demographic studies of a population of each species located within the
469 area of ash fall revealed significant reductions in population density in both taxa during the first
470 summer breeding season following the eruption (E. A. Lacey pers. obs.). Such demographic
471 bottlenecks are typically expected to result in a loss of genetic variation (England et al. 2003).
472 However, contrary to this expectation, we found no evidence of a post-eruption decrease in
473 variation at the mitochondrial cytochrome b locus or multiple microsatellite loci in either *C.*
474 *sociabilis* or *C. haigi*. Instead, our analyses revealed that observed heterozygosity at multiple
475 microsatellite loci was greater in *C. sociabilis*, but not in *C. haigi*, after the 2011 eruption. This
476 apparent post-eruption increase in genetic variation in *C. sociabilis* is intriguing, particularly
477 given the overall low levels of genetic diversity in mitochondrial genes reported for both modern
478 pre-eruption populations of this species and historical populations dating back to 3,000-5,000
479 years before present (Lacey 2001; Hadly et al. 2003; Chan et al. 2005). Collectively, these
480 analyses concur in suggesting that *C. sociabilis* appears to have experienced repeated
481 demographic bottlenecks during its history. The 2011 eruption of the Puyehue-Cordón Caulle

482 complex, however, did not result in an immediate loss of microsatellite genetic diversity but
483 instead led to an unexpected increase in local genetic diversity.

484 *Power of analyses to detect genetic differentiation.*— Our analyses were based on data from the
485 mitochondrial cytochrome b locus and multiple microsatellite loci. While we expected the 2011
486 eruption to have caused decreases in genetic variation, the failure to detect post-eruption changes
487 in cytochrome b in *C. sociabilis* is unsurprising given that only 1 cytochrome b haplotype was
488 present in the focal population of this species prior to the eruption and that variation at this locus
489 appears to have been limited to this single haplotype for at least the past 1,000 years, meaning
490 that there could be no further reduction in genetic variation at this locus (Hadly et al. 2003).

491 Similarly, we find no evidence of a post-eruption decrease in genetic variation in the focal
492 population of *C. haigi*, although such a decrease was possible due to the multiple cytochrome b
493 haplotypes present before the eruption. The ability to infer demographic events from changes in
494 variation at a single locus may be limited, however, as not all portions of the genome are equally
495 capable of capturing the effects of recent demographic events (Matocq and Villablanca 2001;
496 Kilian et al. 2007). The limitations of examining a single mitochondrial gene, combined with a
497 smaller number of individuals sampled in *C. haigi* ($n = 17$), may explain the lack of a signal in
498 cytochrome b for a post-eruption population decline in this species.

499 In addition, although our microsatellite analyses are based on a relatively limited number
500 of loci, the results of our POWSIM analyses indicated that these data should have provided
501 sufficient statistical power to detect pre- and post-eruption differences in genetic diversity as well
502 as differences between the focal and ancillary populations of the same species. An equivalent
503 number of loci have been used to document population differentiation in other taxa (Hale et al.
504 2001; Laikre et al. 2005; Ryman and Palm 2006). Similarly, the number of individuals per

505 population sampled should have been sufficient to provide reasonable estimates of genetic
506 variation, given that a threshold of 25-30 individuals (and 15-20 individuals for populations with
507 high polymorphism) has been suggested to be adequate for quantifying genetic diversity using
508 microsatellites (Hale et al. 2012). These lines of evidence, combined with our findings that
509 multiple loci were concordant in suggesting an increase in post-eruption diversity in *C. sociabilis*,
510 lead us to conclude that this outcome is robust and that our results reflect overall trends in genetic
511 variation in our study species.

512 *Evidence of bottlenecks and past demographic change.*— Our analyses of microsatellite data
513 revealed interspecific differences in the signals for past bottlenecks. Both analyses of
514 heterozygosity excess and *M*-ratio values consistently indicated a past bottleneck in the focal
515 population of *C. sociabilis*. In contrast, no such evidence of past reductions in population size
516 was obtained for the focal population of *C. haigi*. Although this outcome may at first seem
517 incompatible with the documented decrease in population density in both species and the post-
518 eruption increase in microsatellite variation detected for *C. sociabilis*, these tests are best able to
519 detect bottlenecks occurring between 10 and as far as 50 generations ago (Peery et al. 2012). As
520 such, genetic signals of reductions in population size may not become apparent immediately
521 following such an event (Peery et al. 2012; Hoban et al. 2013), suggesting that our evidence for
522 bottlenecks in *C. sociabilis* may reflect demographic changes occurring prior to this study.
523 Indeed, such bottleneck tests have been demonstrated to be most likely to detect ancient
524 bottlenecks resulting in moderate to severe declines in population size (Girod et al. 2011).
525 Consistent with this, *M*-ratio tests of microsatellite data from the same focal population during the
526 pre-eruption period have shown signals of such a bottleneck (Lacey 2001), and paleogenetic data
527 along with Bayesian modeling have provided evidence for a severe bottleneck within the past

528 3,000 years (Hadly et al. 2003; Chan et al. 2006). Thus, it seems likely that the evidence of
529 reductions in population size in *C. sociabilis* reported here reflect older demographic events and
530 not the impacts of the 2011 volcanic eruption.

531 *Interspecific differences in genetic response.*— Several factors may have contributed to the
532 apparent interspecific differences in genetic response to the 2011 eruption reported here.

533 Possibilities include a difference in the deposition of ash between the focal populations of the
534 study species. Although these populations are located immediately across the Limay River from
535 each other, ash depth was greater at the *C. sociabilis* study site (E. A. Lacey pers. obs.). As a
536 result, the consequences of the eruption may have been more severe for *C. sociabilis*. At the same
537 time, the 2 species differ markedly with respect to behavior and demography (Lacey et al. 1997,
538 1998; Lacey and Wieczorek 2004) and these differences may have contributed to the differential
539 genetic responses reported here. In particular, *C. sociabilis* is group living, with multiple closely
540 related adult females sharing the same burrow system and rearing their young communally
541 (Lacey et al. 1997; Izquierdo and Lacey 2008). Groups form due to natal philopatry by females
542 and although all males disperse from their natal burrows, movement of these animals is often
543 within the same local population (Lacey and Wieczorek 2004). In contrast, *C. haigi* is solitary,
544 with individuals of both sexes dispersing from their natal burrow (Lacey et al. 1998). These
545 differences in social behavior and associated dispersal patterns suggest that migration among
546 populations is typically more common in *C. haigi*, which likely contributed to the substantially
547 greater pre-eruption levels of microsatellite diversity in the focal population of this species
548 (Lacey 2001). Together, this background of greater pre-eruption genetic diversity and the
549 presumably higher rates of migration and gene flow among local populations may have served to

550 minimize the impacts of the 2011 eruption on microsatellite diversity in *C. haigi* as compared to
551 *C. sociabilis*.
552 *Demography and increased post-eruption genetic variation.*— The most striking result revealed
553 by our analyses – the apparently greater post-eruption microsatellite variation in *C. sociabilis* –
554 may reflect changes in multiple demographic processes, including drift, mutation, and selection.
555 Each of these processes could influence genetic diversity following a population decline. Our
556 analyses, however, indicate that these factors are unlikely to explain the observed post-eruption
557 change in genetic diversity in *C. sociabilis*. First, our Bayesian modeling demonstrates that our
558 microsatellite results do not fit the expected window of change given random genetic drift (and
559 absent selection, mutation, and migration) after such a demographic bottleneck. In addition, such
560 modeling suggests that a much larger population size than is observed empirically is required to
561 maintain the level of genetic variation observed. Second, these analyses indicate that the mutation
562 rates required to produce the observed change in variation exceed empirical limits for
563 microsatellite loci (Mapelli et al. 2012). Although increased mutation rates have been reported
564 following some catastrophic environmental events, these reports appear to be limited to
565 environmental changes involving known mutagens (e.g., radiation from the Chernobyl nuclear
566 disaster; Dubrova et al. 1996; Ellegren et al. 1997). Third, microsatellites are considered
567 putatively neutral (Li et al. 2002), and given the timescales and putatively random mortality
568 caused by the bottleneck it is unlikely that balancing selection plays a major role following the
569 2011 eruption. Thus, we suggest that genetic drift, mutation, and selection in the microsatellite
570 loci cannot individually account for the post-eruption changes in genetic diversity in *C. sociabilis*.

571 Instead, the increase in microsatellite variation reported here for *C. sociabilis* is most
572 consistent with a scenario of enhanced post-eruption migration and gene flow. Nearby

573 populations of this species may possess different genetic variants, and migration (and thus gene
574 flow) from such populations could have contributed to the observed greater post-eruption genetic
575 diversity in the focal study population. This hypothesis is supported by our demographic
576 modeling, which indicates that the observed increase in variation was unlikely in the absence of
577 migration. Increasing simulated migration rates among nearby populations produced a predicted
578 increase in variation similar to that observed in our empirical data set. Although we do not have
579 the power to ascertain the precise rate of migration that occurred after the 2011 eruption given
580 that effective population sizes for the neighboring populations are unknown, our modeling
581 demonstrates that even small amounts of migration from nearby populations could lead to the
582 levels of increased genetic diversity in the focal population. As such, our results suggest an
583 increase in gene flow from other *C. sociabilis* populations following the 2011 eruption. The
584 ancillary locations sampled (Fig. 1) are all located in close proximity (<1 km) to the focal *C.*
585 *sociabilis* population, with no apparent geographic barriers between populations. Our estimates
586 for the range of migration rates needed to generate the observed post-eruption change in genetic
587 variation are biologically plausible for *C. sociabilis* (Lacey and Wieczorek 2004), and dispersal
588 events occurring over 1-2 km have been detected for this species (Lacey pers. comm.). Such
589 movements are also consistent with patterns of dispersal, migration, and gene flow found in other
590 ctenomyids (Fernández-Stolz et al. 2007; Lopes and de Freitas 2012; Roratto et al. 2015).
591 Additionally, this hypothesis is also consistent with the increased percentage of unmarked
592 (potentially immigrant) females captured in the focal study population of *C. sociabilis* during the
593 breeding season following the eruption (E. A. Lacey pers. obs.). Thus, based on demographic
594 models and empirical data, post-eruption migration and gene flow among local populations

595 appear to provide the most logical explanation for the observed increase in microsatellite genetic
596 diversity in *C. sociabilis*.

597 *Implications for studies of environmental catastrophes.*— Our findings suggest that short-term
598 migration among local populations can play an important role in determining levels of genetic
599 diversity immediately following catastrophic environmental events. In particular, such migration
600 may serve to mitigate expected declines in genetic diversity associated with reductions in
601 effective population size. Increases in local genetic diversity after reductions in population size
602 have been documented for other species, including montane voles (*Microtus montanus*— Hadly
603 et al. 2004) and artesian spring snails (*Fonscochlea accepta*— Wilmer et al. 2011); enhanced
604 migration and gene flow have been suggested as the most likely explanation for these findings.
605 Thus, while catastrophic environmental events are typically expected to result in population
606 bottlenecks and reductions in genetic diversity, actual responses to such changes may be more
607 complex and result in different genetic outcomes. Determining how a given species will respond
608 to such events is challenging and requires detailed information regarding both the nature and
609 magnitude of the environmental change as well detailed information regarding the demography of
610 the organisms in question. As documented here, the 2011 eruption of the Puyehue-Cordón Caulle
611 complex produced different genetic responses in *C. sociabilis* and *C. haigi*, with post-eruption
612 microsatellite heterozygosity increasing in the former species but not the latter. Over the past few
613 thousand years, however, *C. sociabilis* has experienced a decline in genetic variation that is not
614 evident in *C. haigi* (Chan et al. 2005; Chan and Hadly 2011), thereby underscoring the variable
615 nature of responses to environmental conditions. These differences in response – both between
616 species and time periods – raise intriguing questions regarding how interactions among
617 environmental changes, demography, and existing levels of genetic diversity interact to shape

618 responses to a given catastrophic event. We expect migration and gene flow to be an important
619 part of this equation and thus opportunities to combine detailed demographic information with
620 genetic data should prove important in understanding and predicting response to environmental
621 change.

622

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637 **Supplementary data**

638 **Supplementary data S1.** Haplotypic networks throughout the past 12,000 years for A. 398 base
639 pairs of cytochrome b in *C. sociabilis* and B. 388 base pairs of cytochrome b in *C. haigi*. These
640 network maps represent haplotypic change temporally, with the bottom layers reflecting

641 haplotype networks in the past (from ancient DNA) and the top layers reflecting modern
642 diversity.

643

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793

794 **Figure captions**

795 **Figure 1.** A. Map of South America showing inset of field sites. B. Inset depicts the close
796 proximity of the range of *C. sociabilis* range (shaded) to the Puyehue-Cordón Caulle volcanic
797 complex. The range of the parapatric congener *C. haigi* is located across the Limay River. The
798 dashed-dotted line shows approximate cone of tephra from the 2011 eruption. Focal and ancillary
799 populations of *C. sociabilis* (circles) and *C. haigi* (squares) are also indicated. Closed shapes
800 indicate focal populations of *C. sociabilis* (closed circle, at Estancia Rincon Grande) and *C. haigi*
801 (closed square, at Estancia San Ramon). Non-focal populations surround the focal populations

802 and are represented by open circles (*C. sociabilis*) and open squares (*C. haigi*). The localities of
803 paleogenetic samples at Cueva Traful (CT) and Estancia Nahuel Huapi (ENH) are indicated with
804 diamonds.

805
806 **Figure 2.** Haplotype networks for cytochrome b in the focal populations of *C. sociabilis* (on left;
807 900 bp) and *C. haigi* (on right; 600 bp). Each circle represents a unique haplotype, with the size
808 of the circle indicating the relative frequency of that haplotype (number of individuals provided
809 in the center of each circle). Nodes represent the number of base pair differences between
810 haplotypes. The bottom layers (dark gray) for both species depict the haplotype network prior to
811 the 2011 eruption (1993-1998) while the top layers (light gray) depicts the network following the
812 2011 eruption (2012-2013). Solid lines spanning the 2 layers connect any haplotypes that are
813 found in both time periods. Open circles show “missing” haplotypes; these are haplotypes found
814 at other times but not present in the designated time period.

815
816 **Figure 3.** Microsatellite variation in A. *C. sociabilis* and B. *C. haigi*. Graphs depict observed
817 heterozygosity at 5 identified microsatellite loci before (pre-eruption population: 1993-1998) and
818 after (post-eruption population: 2011-2013) the June 2011 volcanic eruption; locus names are
819 abbreviated using Soc followed by the locus number. ** indicates significance at $P < 0.01$ level,
820 Fisher’s exact test

821
822 **Figure 4.** Posterior distributions (in gray) for A. current, post-eruption effective population size
823 and B. mutation rate. Both distributions were generated using an approximate Bayesian
824 computation framework from 1,000,000 simulations run by Bayesian serial coalescent modeling.

825 In each panel, the dashed line indicates the uniform distribution used as a prior distribution. Four
826 mean summary statistics from the empirical data were used: observed heterozygosity before and
827 after the 2011 eruption and allelic variance before and after the 2011 eruption. Posterior
828 distributions for both N_e and mutation rate display sharp peaks for MLE values, indicating high
829 confidence in these values.

830
831 **Figure 5.** Posterior distributions for migration from the collective, pooled ancillary populations
832 of *C. sociabilis* populations into the focal population of this species, as represented forward in
833 time (all simulations were run backward in time but are presented here as forward in time for
834 clarity). The prior for this migration was set as a uniform distribution from 0 to 1 in all
835 simulations. Migration rate is represented on the x-axis, with the scale from 0 to 1 in each graph
836 (see line on bottom left presenting a representative x-axis). Density is depicted on the y-axis
837 (scale varies) and reflect relative probability. Sharper peaks indicate better fit for each model.

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844 **Tables**

845 **Table 1.** Summary of *C. sociabilis* and *C. haigi* samples analyzed by time period and population; totals include 20 pre-eruption
 846 samples from each focal population previously genotyped by Lacey (2001). A. Number of individuals sampled and used in
 847 microsatellite and cytochrome b analyses. B. Age distribution of samples included in each population for microsatellite analyses.
 848 Samples from ancillary populations were not collected after the eruption due to the difficulty of accessing these sites under post-
 849 eruption conditions.

850 A. Number of individuals sampled

	<i>C. sociabilis</i> populations		<i>C. haigi</i> populations	
	Focal	Ancillary	Focal	Ancillary
853 Pre-eruption (1993-1998)				
854 <i>Total samples</i>	74	42	65	25
855 <i>Samples included for microsat analyses</i>	74	42	60	25
856 <i>Samples included for cyt b analyses</i>	14	0	13	0
857 Post-eruption (2011-2013)				
858 <i>Total samples</i>	31	0	17	0
859 <i>Samples included for microsat analyses</i>	31	0	17	0
860 <i>Samples included for cyt b analyses</i>	31	0	17	0

861

862

863 B. Age distribution of individuals sampled

	<i>C. sociabilis</i> populations		<i>C. haigi</i> populations	
	Focal	Ancillary	Focal	Ancillary
866 Pre-eruption (1993-1998)				
867 <i>Adults</i>	57	4	38	23
868 <i>Juveniles</i>	17	38	22	2
869 Post-eruption (2011-2013)				
870 <i>Adults</i>	5	0	10	0
871 <i>Juveniles</i>	26	0	7	0

872 **Table 2.** Summary of microsatellite variation in the focal population of each study species. Results are shown for Hardy-
873 Weinberg probability tests, F_{IS} values (Weir and Cockerham estimate), gene diversity, and expected heterozygosity across the 5
874 microsatellites for A. the pre and post-eruption focal population of *C. sociabilis* and B. the pre- and post-eruption focal
875 population of *C. haigi*

876

877 A. Microsatellite variation in *C. sociabilis*

878 Locus	878 HWE p -value		878 F_{IS} estimates		878 Gene diversity		878 Heterozygosity (H_E)		878 Allelic richness (# alleles)	
	879 pre	879 post	879 pre	879 post	879 pre	879 post	879 pre	879 post	879 pre	879 post
881 Sociabilis 1	0.384	0.291	0.100	-0.261	0.150	0.344	0.150	0.345	1.953 (2)	2.000 (2)
882 Sociabilis 4	0.000	0.234	0.547	-0.429	0.245	0.438	0.244	0.444	2.389 (3)	2.000 (2)
883 Sociabilis 5	0.660	0.636	0.041	-0.152	0.261	0.389	0.261	0.390	1.998 (2)	2.000 (2)
884 Sociabilis 6	1.000	0.359	-0.007	-0.246	0.298	0.432	0.298	0.434	2.228 (3)	2.000 (2)
885 Sociabilis 7	1.000	0.060	0.028	-0.410	0.319	0.422	0.319	0.425	2.000 (3)	2.000 (2)

886

887 B. Microsatellite variation in *C. haigi*

888 Locus	888 HWE p -value		888 F_{IS} estimates		888 Gene diversity		888 Heterozygosity (H_E)		888 Allelic richness (# alleles)	
	889 pre	889 post	889 pre	889 post	889 pre	889 post	889 pre	889 post	889 pre	889 post
891 Sociabilis 1	0.337	0.436	0.017	0.020	0.821	0.829	0.821	0.829	7.213 (9)	7.000 (7)
892 Sociabilis 4	0.495	0.820	-0.016	0.033	0.627	0.608	0.627	0.608	4.492 (5)	4.882 (5)
893 Sociabilis 5	0.402	0.805	0.070	-0.044	0.723	0.846	0.722	0.847	6.792 (10)	6.941 (7)
894 Sociabilis 6	0.476	0.013	-0.015	0.078	0.752	0.827	0.752	0.825	8.561 (13)	7.879 (9)
895 Sociabilis 7	0.581	0.779	0.068	-0.062	0.855	0.831	0.854	0.832	9.240 (12)	8.763 (9)

896 **Table 3.** Tests for population bottlenecks in *C. sociabilis* and *C. haigi*. Comparisons of observed *M*-ratios to critical values of *M*
897 for A. the post-2011 eruption focal population of *C. sociabilis* and B. the post-2011 eruption focal population of *C. haigi*. Critical
898 *M*-values were computed with a range of values of θ ($4N_e\mu$) ranging from 1.2 to 12, which factors in an effective population size
899 of 300 and a range of μ from 10^{-3} to 10^{-4} . An asterisk (*) indicates significance at the $P < 0.05$ level because critical *M* values
900 represent the point where 95% of *M*-ratios at equilibrium will be above that critical value. A dash (-) represents non-significant
901 values. Critical *M* values are not repeated after the first row since they are the same for each locus, given identical number of
902 samples, loci, and parameters in the model.

903

904 A. Population bottleneck tests for *C. sociabilis*

905 Locus	<i>M</i> -ratio	$\theta=1.2$	$\theta=5$	$\theta=10$	$\theta=12$
906 Sociabilis 1	0.286	0.717*	0.660*	0.639*	0.637*
907 Sociabilis 4	0.500	*	*	*	*
908 Sociabilis 5	0.222	*	*	*	*
909 Sociabilis 6	0.250	*	*	*	*
910 Sociabilis 7	0.333	*	*	*	*

912

913

914 B. Population bottleneck tests for *C. haigi*

915 Locus	<i>M</i> -ratio	$\theta=1.2$	$\theta=5$	$\theta=10$	$\theta=12$
916 Sociabilis 1	1.000	0.713	0.634	0.599	0.583
917 Sociabilis 4	1.000	-	-	-	-
918 Sociabilis 5	0.538	*	*	*	*
919 Sociabilis 6	0.727	-	-	-	-
920 Sociabilis 7	0.750	-	-	-	-

922